

π IN THE SKY⁴

You don't have to be a NASA rocket scientist to solve this Martian crater mystery.
All you need is a little pi!

Discover more “ π in the sky” math problems at:
jpl.nasa.gov/edu/nasapidaychallenge

CRATER CURIOSITY

Craters form when an object hits the surface of a planet or other body. The impact creates a round impression surrounded by material, called ejecta, that gets blasted out of the crater. Scientists study ejecta because it contains clues about what's below a planet's surface. When an object hits Mars at an angle under 20 degrees, the crater is less circular and the ejecta settles in a butterfly shape. Some areas around the crater contain no blast material. Finding craters that formed this way can help scientists understand how meteor impacts change the surface of a planet. To do this, they measure a crater's circularity ratio. If the ratio is less than 0.925, it suggests that an object impacted at an angle under 20 degrees and created a butterfly pattern.

Use the circularity ratio formula, $\frac{4\pi A}{p^2}$, to determine which of these craters would have the butterfly ejecta pattern.

LEARN MORE ABOUT CRATERS
bit.ly/mars

UNNAMED CRATER  $A = 32 \text{ km}^2$
 $p = 21 \text{ km}$

AVEIRO CRATER  $A \text{ (area)} = 67 \text{ km}^2$
 $p \text{ (perimeter)} = 30 \text{ km}$

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You don't have to be a NASA rocket scientist to measure the size of the moon's shadow during a total solar eclipse. All you need is a little π !

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EPIC ECLIPSE

When sunlight hits the moon, a cone-shaped shadow is created. During the total solar eclipse on August 21, 2017, the distance from the center of the moon to the center of Earth will be 372,027 km. On that day, if the moon's shadow were not intersected by the surface of Earth, it would extend 377,700 km from the moon to its vertex.

Viewers on Earth who want to witness the eclipse will have to be at a location inside this shadow as it passes over Earth to see the eclipse at totality. **What is the approximate surface area of Earth that will be covered by the disc of the moon's shadow at any one time during the eclipse?**

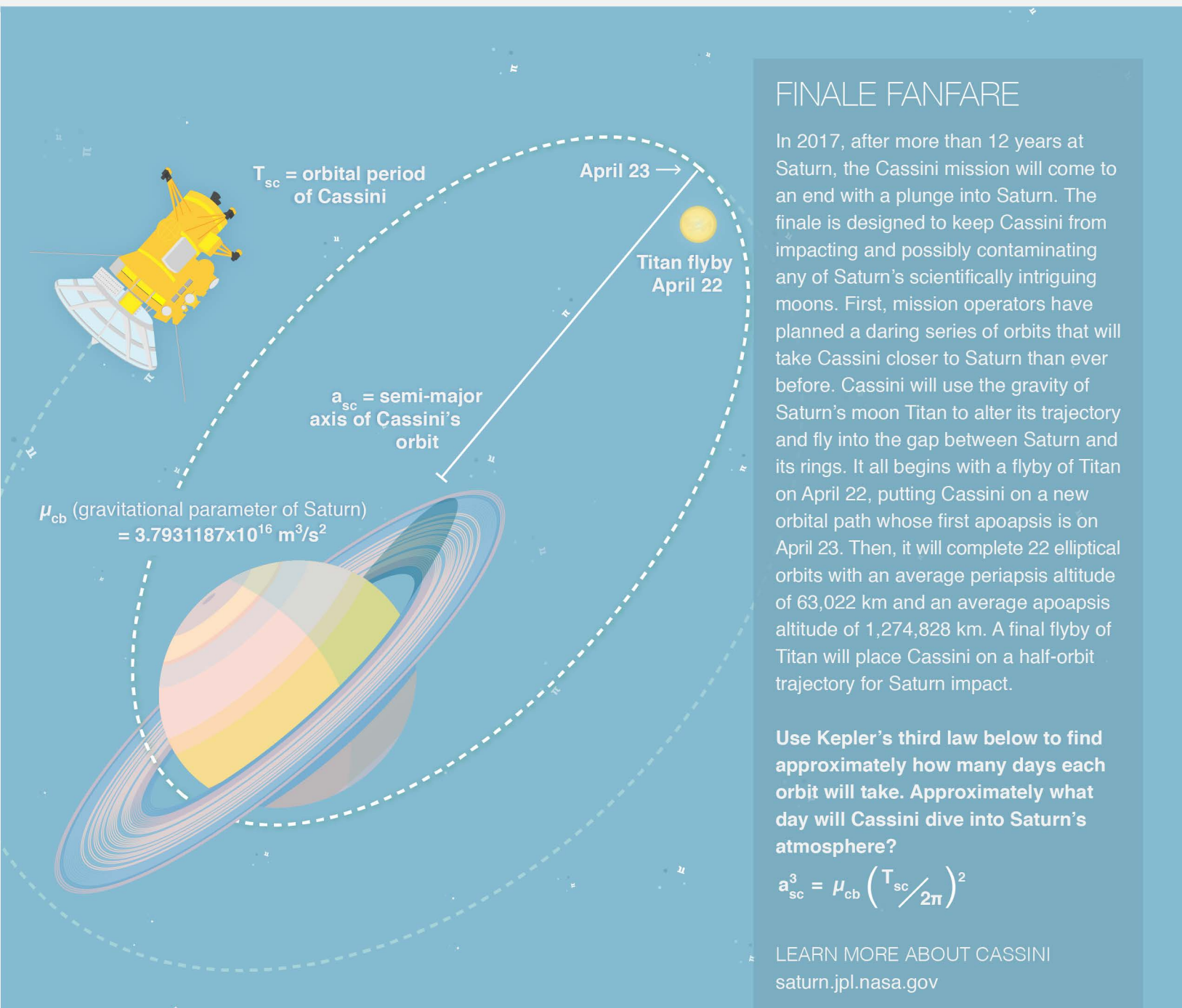
LEARN MORE ABOUT THE ECLIPSE
eclipse2017.nasa.gov



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You don't have to be a NASA rocket scientist to plot the trajectory for Cassini's daring orbit around Saturn. All you need is a little π !

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π IN THE SKY⁴

You don't have to be a NASA rocket scientist to discover potentially habitable worlds outside our solar system. All you need is a little pi!

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HABITABLE HUNT

Scientists can learn a lot about planets beyond our solar system by studying their stars. They can calculate an exoplanet's orbital period by measuring how often its star dims as the planet passes by. They can even find potentially habitable worlds with a few key details. The star's temperature and luminosity, which are related to its mass, define its habitable zone, the area where liquid water can exist. And the bond albedo, or percentage of light reflected by the exoplanet, helps estimate its temperature. Scientists recently discovered seven Earth-like planets orbiting the star TRAPPIST-1. **Given TRAPPIST-1's measurements, what are the inner and outer radii (r), in AU, of its habitable zone?** Use the formula below.

$$r = \sqrt{\frac{(1-A)L}{16\pi\sigma T^4}}$$

Given the orbital periods (T_p), for TRAPPIST-1's planets, which are in the habitable zone? Use Kepler's third law to find the semi-major axis of each planet's orbit (a_p).

$$a_p^3 = \mu_{cb} \left(\frac{T_p}{2\pi} \right)^2$$

LEARN MORE ABOUT EXOPLANETS
exoplanets.nasa.gov

TRAPPIST-1 SYSTEM

L (star luminosity) = 2.0097×10^{23} watts

μ_{cb} (star gravitational parameter) =
 $1.06198 \times 10^{19} \text{ m}^3/\text{s}^2$

σ (Stefan-Boltzmann constant) =
 $5.67 \times 10^{-8} \text{ Wm}^2\text{K}^{-4}$

T (planetary temperature) = 192-295 K

A (planetary bond albedo) = 0.3

